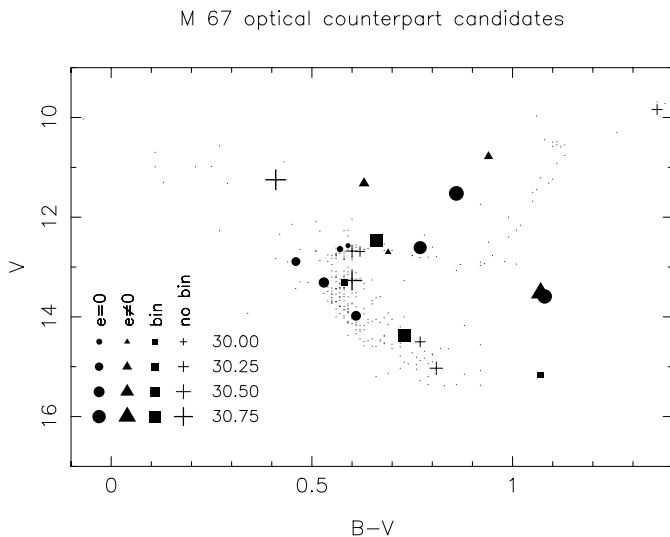


*Erratum***X rays from old open clusters: M 67 and NGC 188**T. Belloni<sup>1,\*</sup>, F. Verbunt<sup>2</sup>, and R.D. Mathieu<sup>3</sup><sup>1</sup> Astronomical Institute “Anton Pannekoek”, University of Amsterdam and Center for High-Energy Astrophysics, Kruislaan 403, 1098 SJ Amsterdam, The Netherlands<sup>2</sup> Astronomical Institute, P.O.Box 80000, 3508 TA Utrecht, The Netherlands<sup>3</sup> Department of Astronomy, University of Wisconsin, Madison, WI 53706, USAAstron. Astrophys. (1998) **339**, 431–439

The correct  $B - V$  colour of Sanders 628, the suggested counterpart of X-ray source no. 35 in Table 1 of Belloni et al. (1998) is  $B - V = 0.77$ . The location of the counterpart thus is close to the main sequence. The suggestion in Table 3 of Belloni et al. (1998) that the source is a cataclysmic variable is no longer viable; it is more likely to be a RS CVn type system.

The corrected version of Fig. 1 is shown.



**Fig. 1.** HR diagram of M 67 in which the stars detected in X rays are indicated with special symbols. The size of a symbol is proportional to the logarithm of the X-ray luminosity in the 0.1–2.4 keV band in  $\text{erg s}^{-1}$ . The shape of a symbol indicates its spectroscopic binary status. Filled circles and triangles indicate circular binaries and eccentric binaries (i.e. eccentricity larger than  $3\sigma$ ), respectively, squares indicate binaries with unknown period, and crosses indicate objects for which there is no indication of binarity. Note that for the binaries we show the total magnitude and colour.

---

Send offprint requests to: T. Belloni (belloni@merate.mi.astro.it)

\* Present address: Osservatorio astronomico di Brera, Via E. Bianchi 46, 23807, Merate (LC), Italy